

DPP

DAILY PRACTICE PROBLEMS

Class : XIIth
Date :

Solutions

Subject : PHYSICS
DPP No. : 9

Topic :- NUCLEI

1

(b)

$$\text{Here } \frac{N}{N_0} = \left(\frac{1}{2}\right)^n = \left(\frac{1}{2}\right)^{1/3}$$

$$\text{Where } n = \text{Number of half lives} = \frac{1}{3}$$

$$\Rightarrow \frac{N}{N_0} = \frac{1}{1.26} \Rightarrow \frac{N_U}{N_{Pb} + N_U} = \frac{1}{1.26}$$

$$\Rightarrow N_{Pb} = 0.26 N_U \Rightarrow \frac{N_{Pb}}{N_U} = 0.26$$

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(a)

According to Rydberg's formula

$$\frac{1}{\lambda} = R \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$$

Here, $n_f = 1, n_i = n$

$$\therefore \frac{1}{\lambda} = R \left(\frac{1}{1^2} - \frac{1}{n^2} \right) \Rightarrow \frac{1}{\lambda} = R \left(1 - \frac{1}{n^2} \right) \dots (i)$$

Multiplying equation (i) by λ on both sides,

$$1 = \lambda R \left(1 - \frac{1}{n^2} \right) \Rightarrow \frac{1}{\lambda R} = 1 - \frac{1}{n^2}$$

$$\Rightarrow \frac{1}{n^2} = 1 - \frac{1}{\lambda R} \Rightarrow \frac{1}{n^2} = \frac{\lambda R - 1}{\lambda R} \Rightarrow n = \sqrt{\frac{\lambda R}{\lambda R - 1}}$$

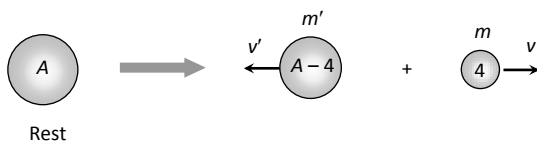
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(c)

Energy of stars is due to the fusion of light hydrogen nuclei into *He*. In this process much energy is released

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(a)



According to conservation of momentum $4v = (A - 4)v'$

$$\Rightarrow v' = \frac{4v}{A - 4}$$

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(c)For third line of Balmer series $n_1 = 2, n_2 = 5$

$$\therefore \frac{1}{\lambda} = RZ^2 \left[\frac{1}{n_1^2} - \frac{1}{n_2^2} \right] \text{ gives } Z^2 = \frac{n_1^2 n_2^2}{(n_2^2 - n_1^2) \lambda R}$$

On putting values $Z = 2$

$$\text{From } E = -\frac{13.6Z^2}{n^2} = \frac{-13.6(2)^2}{(1)^2} = -54.4 \text{ eV}$$

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(d)Using conservation of momentum $P_{\text{daughter}} = P_{\alpha}$

$$\Rightarrow \frac{E_d}{E_{\alpha}} = \frac{m_{\alpha}}{m_d} \Rightarrow E_d = \frac{E_{\alpha} \times m_{\alpha}}{m_d} = \frac{6.7 \times 4}{214} = 0.125 \text{ MeV}$$

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(d) $B.E.$ per nucleon \propto stability

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(a)According to Bohr theory, $mvr = n \frac{h}{2\pi} \Rightarrow v = \frac{nh}{2\pi mr}$

$$\text{and } \frac{mv^2}{r} \propto \frac{k}{r} \Rightarrow \frac{m}{r} \left(\frac{n^2 h^2}{4\pi^2 m^2 r^2} \right) \propto \frac{k}{r} \Rightarrow r_n \propto n$$

$$\text{Kinetic energy } T = \frac{1}{2} m v^2 = \frac{1}{2} m \left(\frac{n^2 h^2}{4\pi^2 m^2 r^2} \right) \Rightarrow T_n \propto \frac{n^2}{r^2}$$

But as $r \propto n$ therefore $T \propto n^0$

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(a)

$$\text{For Lyman series } \nu = RC \left[\frac{1}{1^2} - \frac{1}{n^2} \right]$$

Where $n = 2, 3, 4, \dots$ For the series limit of Lyman series $n = \infty$

$$\therefore \nu_1 = RC \left[\frac{1}{1^2} - \frac{1}{\infty^2} \right] = RC \dots \text{(i)}$$

For the first line of Lyman series, $n = 2$

$$\therefore \nu_2 = RC \left[\frac{1}{1^2} - \frac{1}{2^2} \right] = \frac{3}{4} RC \dots \text{(ii)}$$

$$\text{For Balmer series } \nu = RC \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$$

Where $n = 3, 4, 5 \dots$ For the series limit of Balmer series $n = \infty$

$$\therefore \nu_3 = RC \left[\frac{1}{2^2} - \frac{1}{\infty^2} \right] = \frac{RC}{4} \dots \text{(iii)}$$

From equations (i), (ii) and (iii), we get

$$\nu_1 = \nu_2 + \nu_3 \Rightarrow \nu_1 - \nu_2 = \nu_3$$

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(b)

Positron is the antiparticle of electron

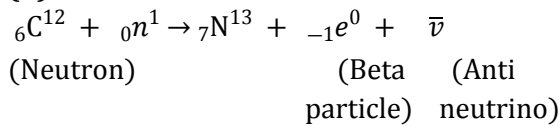
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(d)Nuclides with same atomic number Z but different mass number A are known as isotopesNuclides with same mass number A but different atomic number Z are known as isobarsNuclides with same neutron number $N = (A - Z)$ but different atomic number Z are known as isotones ${}_1H^2$ and ${}_1H^3$ are isotopes

${}^2_2\text{He}^3$ and ${}^1_1\text{H}^3$ are isobars
 ${}^{197}_{79}\text{Au}$ and ${}^{198}_{80}\text{Hg}$ are isotones

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(b)



On equating atomic numbers and atomic masses, the atomic number and atomic mass for resulting nucleus is 7 and 13, which is for nitrogen nucleus.

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(d)

$$E = \Delta mc^2 \Rightarrow E = \frac{0.3}{1000} \times (3 \times 10^8)^2 = 2.7 \times 10^{13} \text{J}$$

$$= \frac{2.7 \times 10^{13}}{3.6 \times 10^6} = 7.5 \times 10^6 \text{kWh}$$

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(d)

The number force is charge independent

No. of nucleons = No. of protons + no. of neutrons = Mass number

All nuclei have masses that are less than the sum of the masses of its constituents. The difference in mass of a nucleus and its constituents is known as mass defect.

Nucleons belong to the family of hadrons while electrons belong to family of leptons

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(a)

Given $N_0\lambda = 5000$, $N\lambda = 1250$

$$N = N_0 e^{-\lambda t} = N_0 e^{-5\lambda}$$

Where λ is decay constant per min.

$$N\lambda = N_0\lambda e^{-5\lambda}$$

$$1250 = N_0\lambda e^{-5\lambda}$$

$$\therefore e^{-5\lambda} = \frac{5000}{1250} = 4$$

$$e^{5\lambda} = 4$$

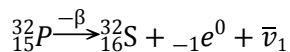
$$5\lambda = 2 \log_e 4$$

$$\lambda = 0.4 \ln 2$$

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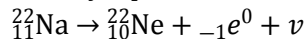
(d)

β^- -emission takes place from a radioactive nucleus as



Where $\bar{\nu}$ is the anti-neutrino.

In β^+ decay a positron is emitted as



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(a)

$$\text{Excitation energy } \Delta E = E_2 - E_1 = 13.6 Z^2 \left[\frac{1}{1^2} - \frac{1}{2^2} \right]$$

$$\Rightarrow 40.8 = 13.6 \times \frac{3}{4} \times Z^2 \Rightarrow Z = 2$$

$$\text{Now required energy to remove the electron from ground state} = \frac{+13.6Z^2}{(1)^2} = 13.6(Z)^2$$

$$= 54.4 \text{ eV}$$

| ANSWER-KEY | | | | | | | | | | |
|------------|----|----|----|----|----|----|----|----|----|----|
| Q. | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 |
| A. | B | B | A | C | A | A | C | D | D | A |
| | | | | | | | | | | |
| Q. | 11 | 12 | 13 | 14 | 15 | 16 | 17 | 18 | 19 | 20 |
| A. | A | B | D | B | D | D | C | A | D | A |
| | | | | | | | | | | |

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